

# Meaning of temperature in different thermostistical ensembles



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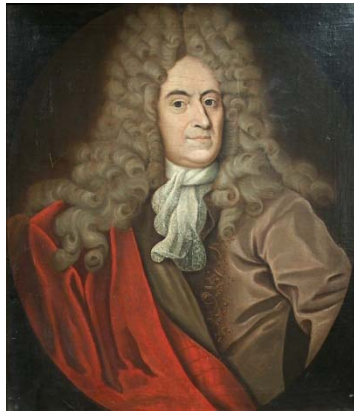
# First thermometer & temperature scales

- 1638: Robert Fludd – air thermometer & scale
- ~1700: linseed oil thermometer by Newton
- 1701: red wine as temperature indicator by Rømer
- 1702: Guillaume Amontons: Absolute zero temperature?
- 1714: mercury and alcohol thermometer by Fahrenheit

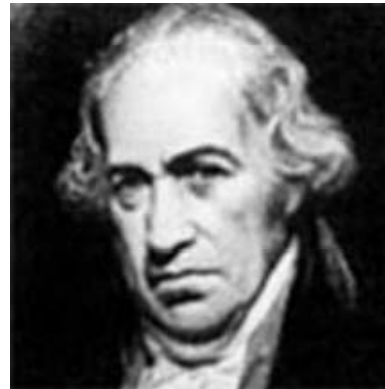
## Definition of temperature scales



**Sir Isaac Newton**  
(1643 – 1727)



**Olaf Christensen  
Rømer**  
(1644 – 1710)



**Daniel Gabriel  
Fahrenheit**  
(1686 – 1736)

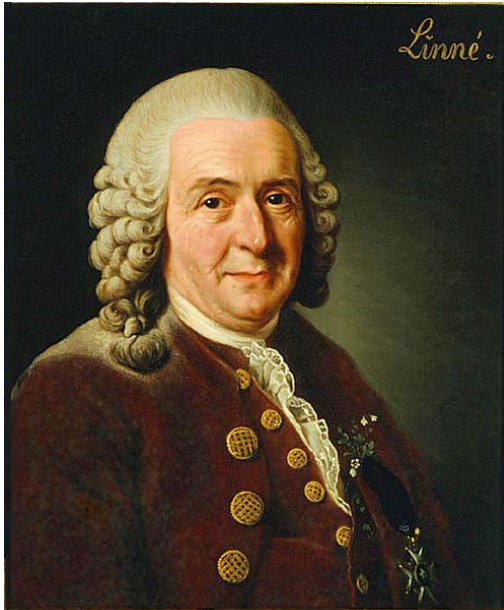


**René Antoine  
Ferchault de  
Réaumur**  
(1683 – 1757)



**Anders Celsius**  
(1701 – 1744)

# Linneaus thermometer

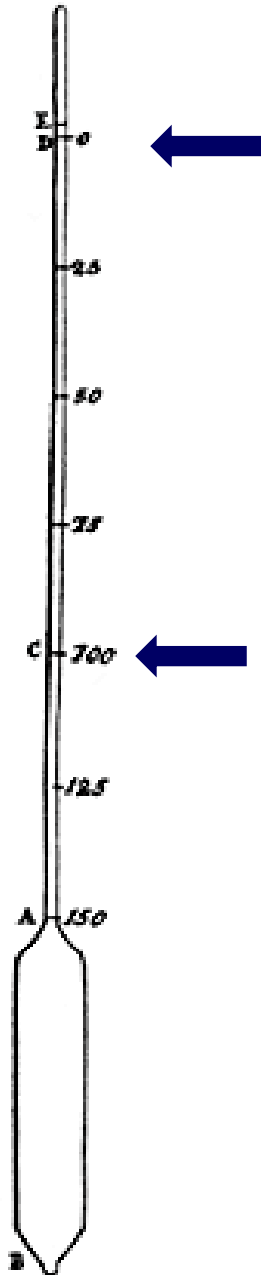


**Carl von Linné**  
(1707 – 1778)

Reversed the Celsius scale

1744: broken on delivery

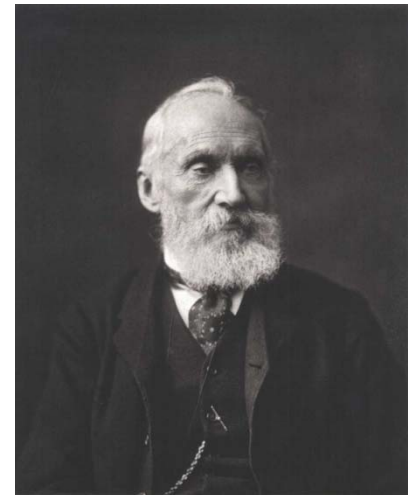
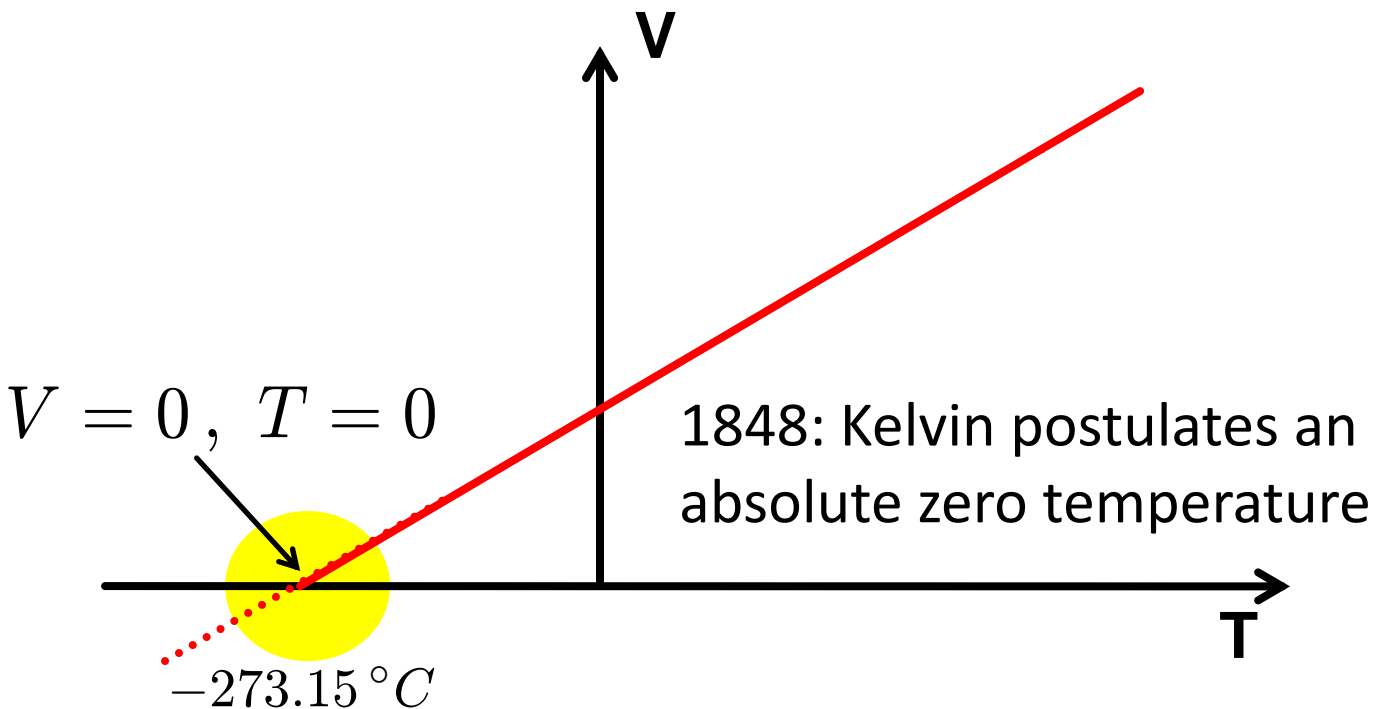
1745: botanical garden in Uppsala



Anders Celsius

# Absolute Zero

Ideal gas law:  $p V = N k_B T$



**William Thomson**  
— Lord Kelvin  
(1824 – 1907)

*It is impossible by any procedure to reduce the temperature of a system to zero in a finite number of operations.*

# *The highest temperature you can see*

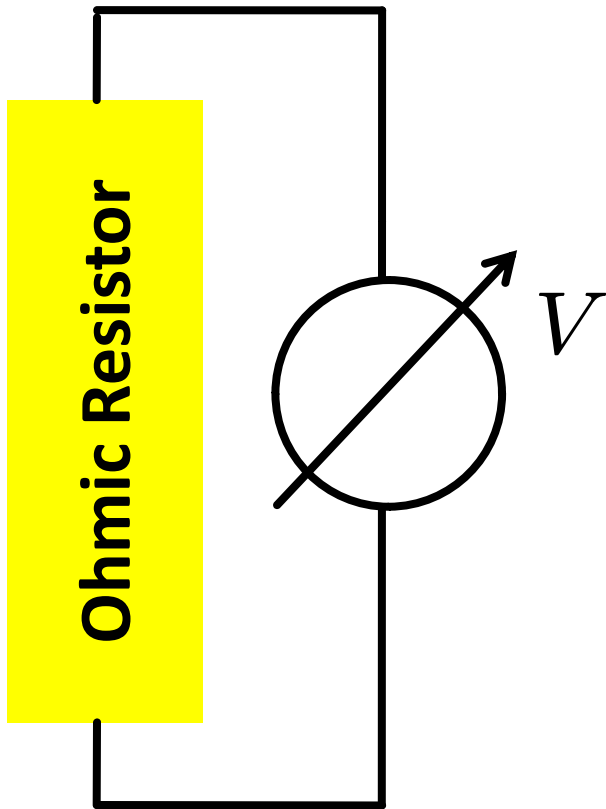


**Lightning:**

**30 000 °C**

Fuse soil or sand into glass

# Noise Thermometer



Johnson – Nyquist noise

$$\text{PSD}_V(\omega) = 2k_B T R$$

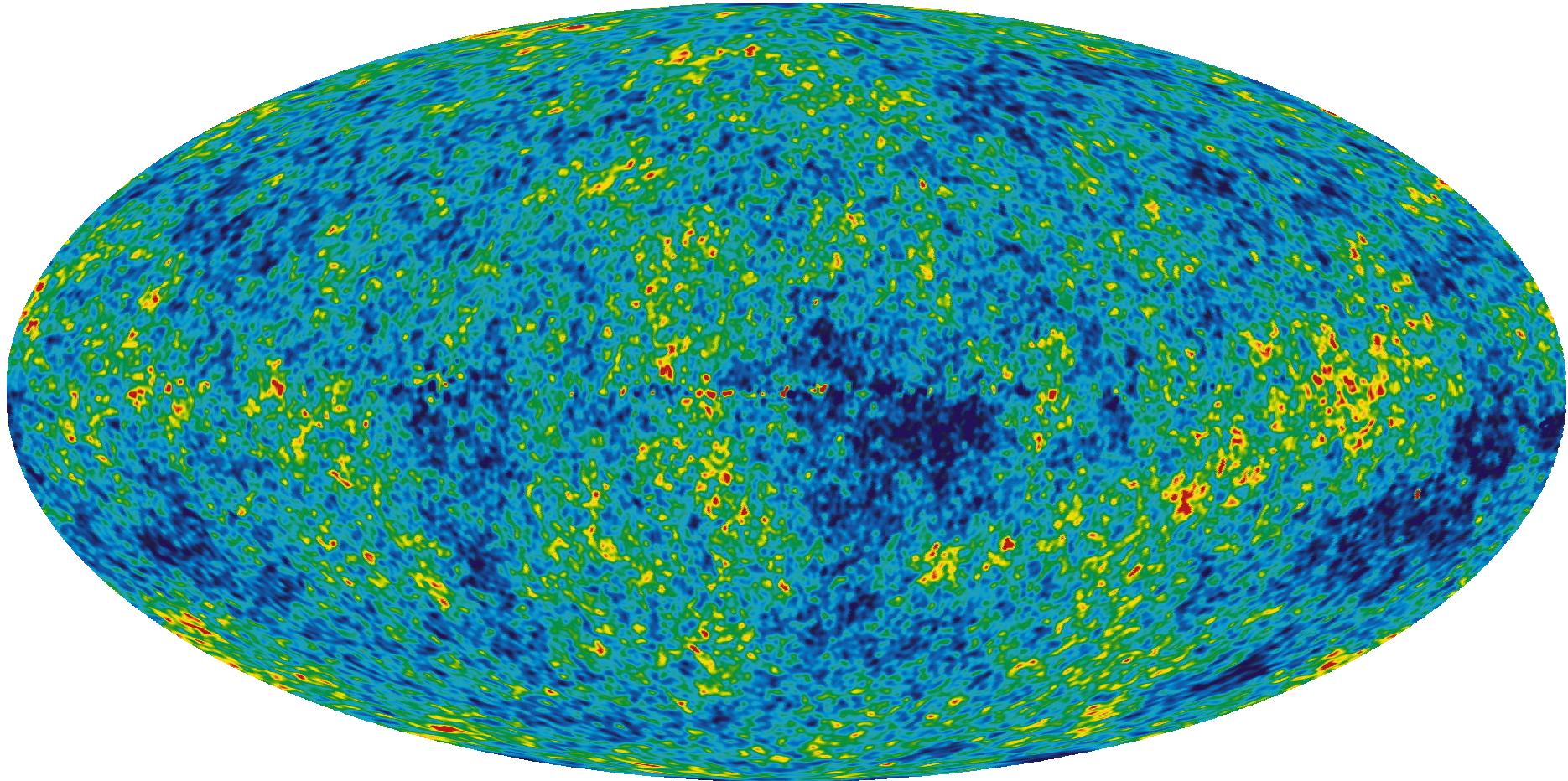
-- classical regime only --

$\text{PSD}_V$  : power spectral density  
of the voltage signal

$k_B$  : Boltzmann constant

$R$  : resistance

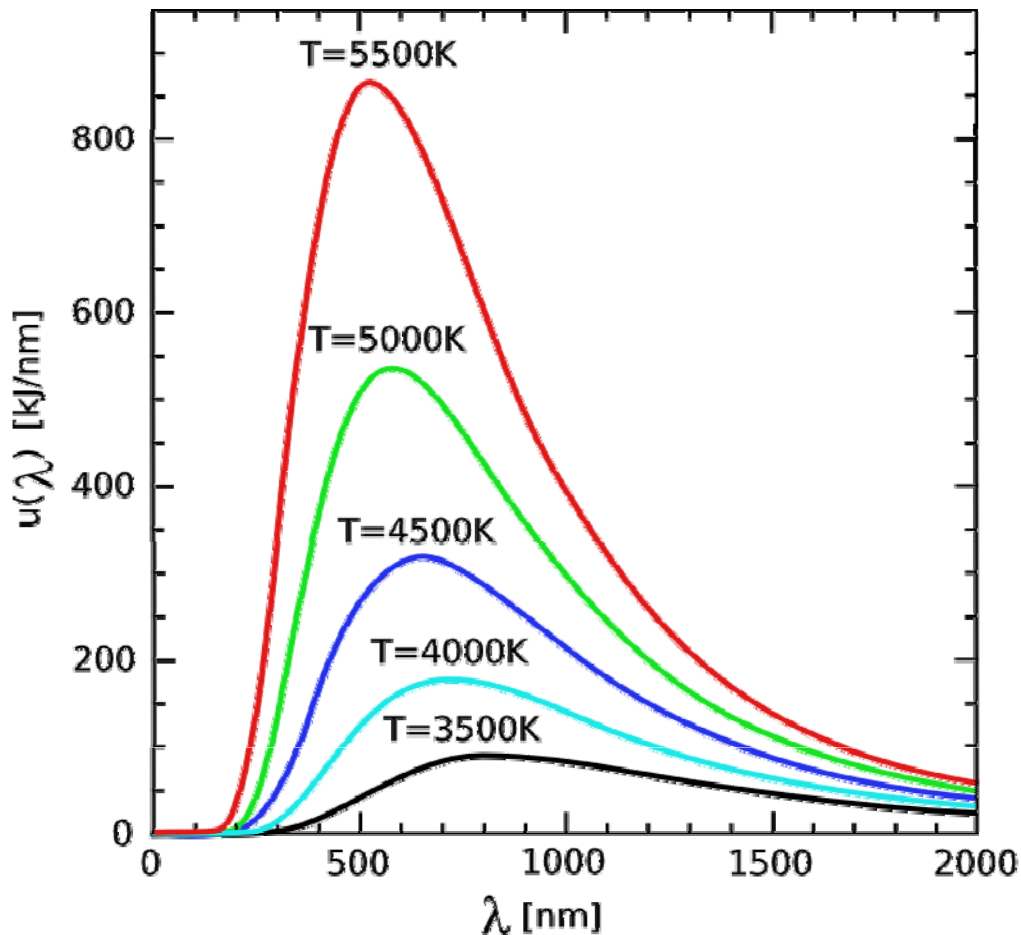
# Cosmic background temperature



$$T = 2.725 \pm \dots \text{ K}$$

# Black body radiation

Planck's law [1901]:  $u(\lambda, T) = \frac{8\pi hc}{\lambda^5} \frac{1}{\exp\left(\frac{hc}{\lambda k_B T}\right) - 1}$



$u(\lambda, T)$  : spectral energy density

$\lambda$  : wavelength

$h$  : Planck constant

$c$  : speed of light

$k_B$  : Boltzmann constant

Stefan – Boltzmann law:

$$E \propto T^4$$

**Thermometer !**



# The famous Laws

## Equilibrium Principle -- minus first Law

*An isolated, macroscopic system which is placed in an arbitrary initial state within a finite fixed volume will attain a unique state of equilibrium.*

## Second Law (Clausius)

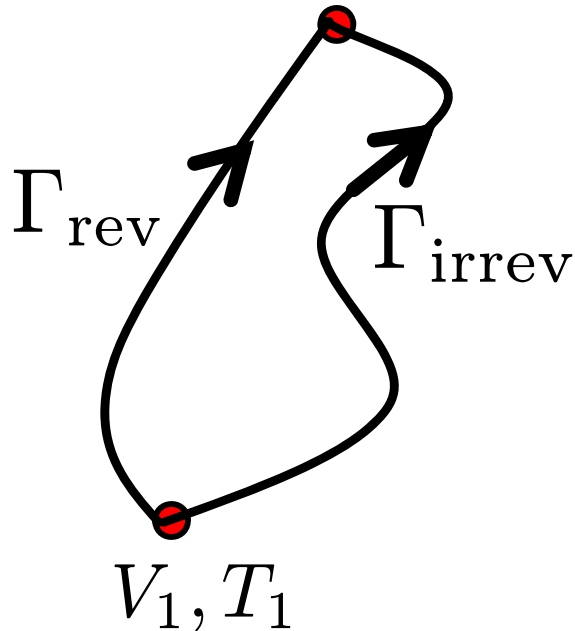
*For a **non-quasi-static process** occurring in a **thermally isolated system**, the entropy change between two equilibrium states is non-negative.*

## Second Law (Kelvin)



# Entropy $S$ – content of *transformation* „Verwandlungswert“

$$dS_{V_2, T_2} = \delta Q^{\text{rev}} / T; \quad \delta Q^{\text{irrev}} < \delta Q^{\text{rev}}$$

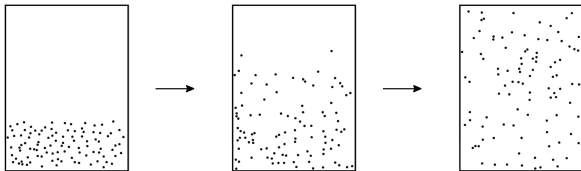


$$\oint_C \frac{\delta Q}{T} \leq 0$$

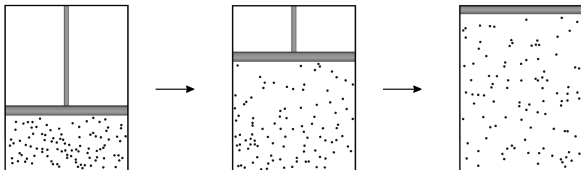
$$C = \Gamma_{\text{rev}}^{-1} + \Gamma_{\text{irrev}}$$

$$\left. \begin{aligned} S(V_2, T_2) - S(V_1, T_1) &\geq \int_{\Gamma_{\text{irrev}}} \frac{\delta Q}{T} \\ S(V_2, T_2) - S(V_1, T_1) &= \int_{\Gamma_{\text{rev}}} \frac{\delta Q^{\text{rev}}}{T} \end{aligned} \right\} \frac{\partial S}{\partial t} \geq 0 \quad \text{NO !}$$

# MINUS FIRST LAW vs. SECOND LAW



**-1st Law**



**2nd Law**

# SECOND LAW

Quote by Sir Arthur Stanley Eddington:

**“If someone points out to you that your pet theory of the universe is in disagreement with Maxwell’s equations – then so much the worse for Maxwell’s equations. If it is found to be contradicted by observation – well, these experimentalists do bungle things sometimes. But if your theory is found to be against the second law of thermodynamics I can give you no hope; there is nothing for it but to collapse in deepest humiliation.”**

Freely translated into German:

**Falls Ihnen jemand zeigt, dass Ihre Lieblingstheorie des Universums nicht mit den Maxwellgleichungen übereinstimmt - Pech für die Maxwellgleichungen. Falls die Beobachtungen ihr widersprechen - nun ja, diese Experimentatoren bauen manchmal Mist. -- Aber wenn Ihre Theorie nicht mit dem zweiten Hauptsatz der Thermodynamik übereinstimmt, dann kann ich Ihnen keine Hoffnung machen; ihr bleibt nichts übrig als in tiefster Schande aufzugeben.**

# Thermodynamic Temperature

$$\delta Q^{\text{rev}} = T dS \leftarrow \text{thermodynamic entropy}$$

$$S = S(E, V, N_1, N_2, \dots; M, P, \dots)$$

$S(E, \dots)$ : (continuous) & differentiable and

**monotonic** function of the internal energy  $E$

$$\left( \frac{\partial S}{\partial E} \right)_{\dots} = \frac{1}{T}$$

**microcanonical ensemble**

# Entropy in Stat. Mech.

$$S = k_B \ln \Omega(E, V, \dots)$$

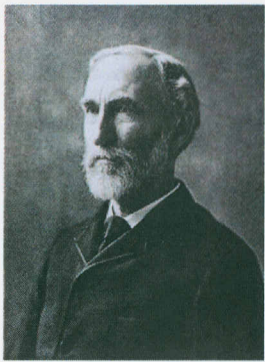
QM:  $\Omega_G(E, V, \dots) = \sum_{0 \leq E_i \leq E} 1$



**classical**

Gibbs:  $\Omega_G = \left( \frac{1}{N! h^{\text{DOF}}} \right) \int d\Gamma \Theta(E - H(\underline{q}, \underline{p}; V, \dots))$

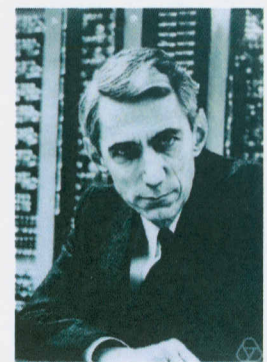
Boltzmann:  $\Omega_B = \epsilon_0 \frac{\partial \Omega_G}{\partial E} \propto \int d\Gamma \delta(E - H(\underline{q}, \underline{p}; V, \dots))$   
density of states



J. W. Gibbs



L. Boltzmann



C. E. Shannon

$$H_G = \int W_N \ln W_N d\Gamma_N$$

$$H_B = N \int W_i \ln W_i d\Gamma_i$$

$$S_G = k_B \ln \Omega_G$$

$$S_B = k_B \ln \left( \frac{\partial \Omega_G}{\partial E} \right) \delta E$$

$$S_S = - \sum_i p_i \log_2 p_i$$

⊕  
ZOO

Renyi  
 Relative  
 Kullback-Leibler  
 v. NEUMANN  
 COLMOGOROV-SINAI  
 FISHER  
 Daroczy-Harvda-Charvat-Tsallis  
 WEHRL  
 Hartley-Chaitin  
 CLAUDIUS  
 CONDITIONAL  
 ETC.



# Entropy in Stat. Mech.

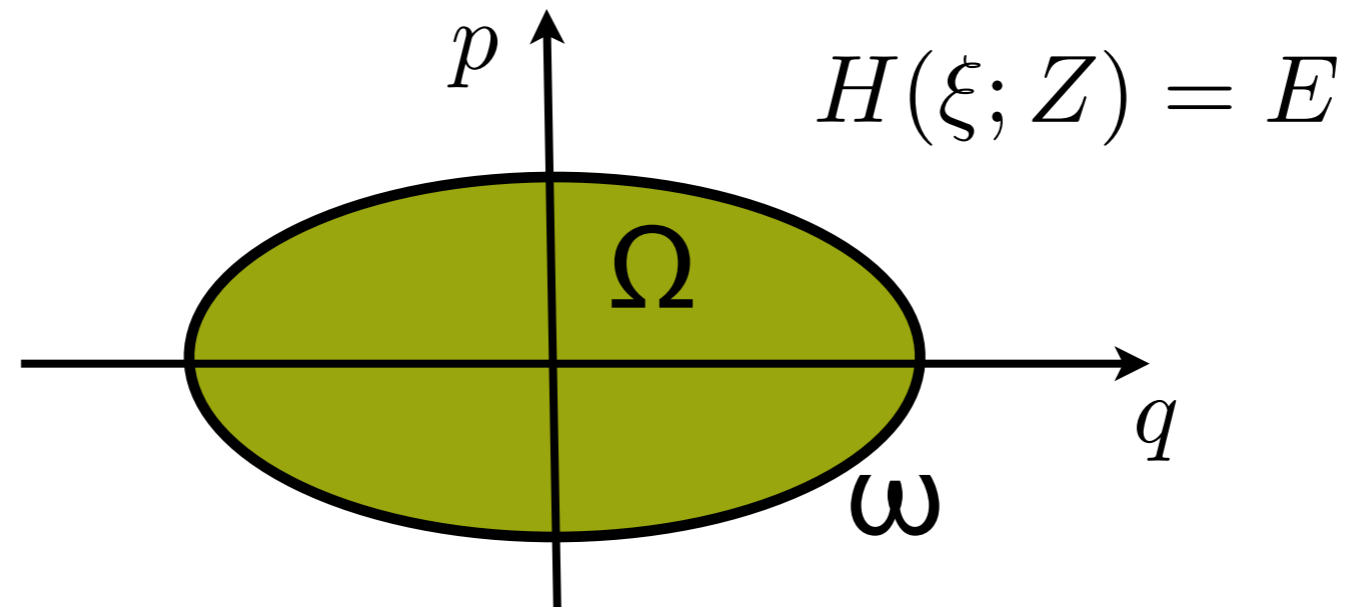
$$S = k_B \ln \Omega(E, V, \dots)$$

**Gibbs:**  $\Omega_G = \left( \frac{1}{N! h^{\text{DOF}}} \right) \int d\Gamma \Theta(E - H(\underline{q}, \underline{p}; V, \dots))$

**Boltzmann:**  $\Omega_B = \epsilon_0 \frac{\partial \Omega_G}{\partial E} \propto \int d\Gamma \delta(E - H(\underline{q}, \underline{p}; V, \dots))$

density of states

# Microcanonical thermostatics



D-Operator

DoS

$$\rho(\xi|E, Z) = \frac{\delta(E - H)}{\omega}$$

$$\omega(E, Z) = \text{Tr}[\delta(E - H)] \geq 0$$

$$\Omega(E, Z) = \text{Tr}[\Theta(E - H)]$$

IntDoS

## Thermodynamic Entropy ?

$$S_B(E) = \ln(\epsilon \omega)$$

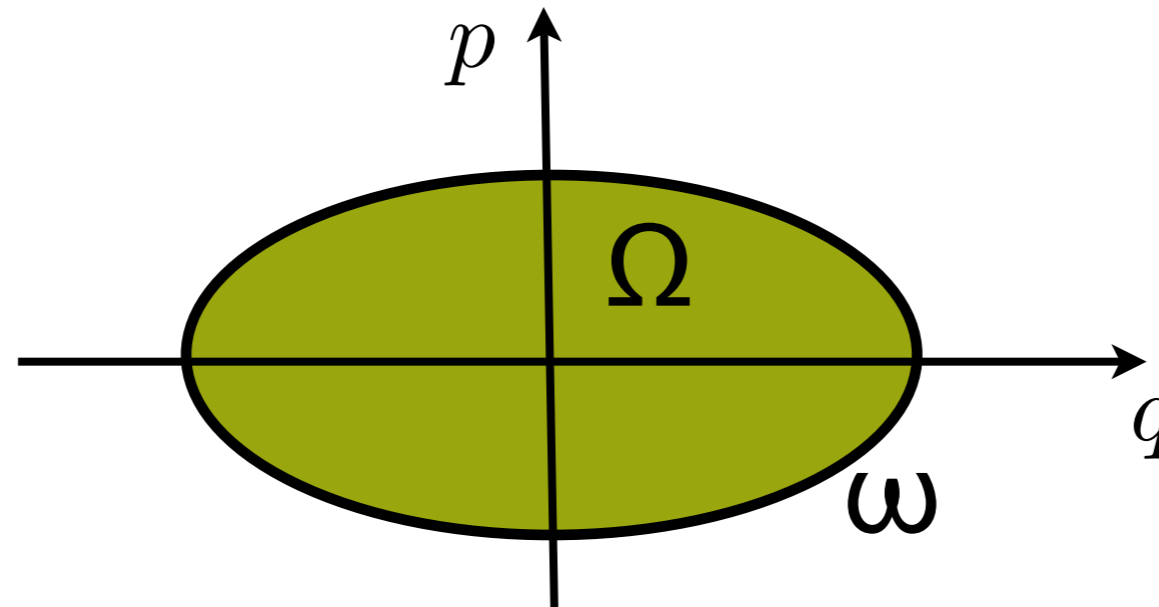
Boltzmann (?)

vs.

$$S_G(E) = \ln \Omega$$

Gibbs (1902), Hertz (1910)

# Boltzmann vs. Gibbs



$$S_B(E) = \ln(\epsilon \omega)$$

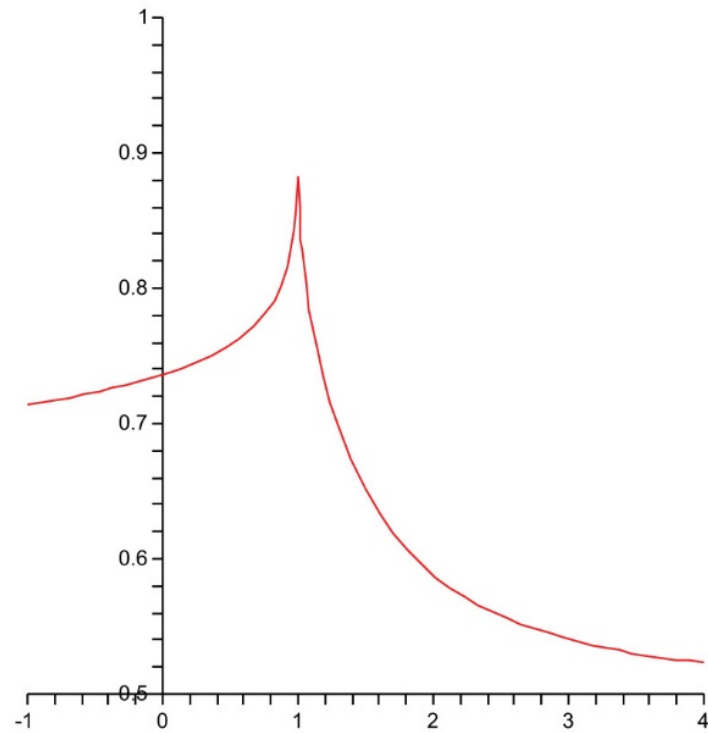
$$S_G(E) = \ln \Omega$$

$$T(E, Z) \equiv \left( \frac{\partial S}{\partial E} \right)^{-1}$$

$$T_B(E) = \frac{\omega}{\nu} \gtrless 0$$

$$T_G(E) = \frac{\Omega}{\omega} \geq 0$$

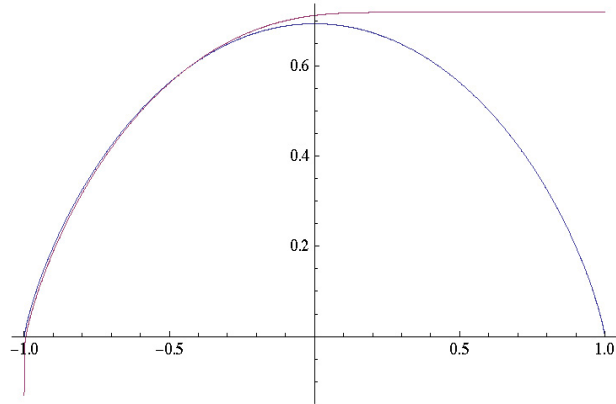
$$\nu(E, Z) = \partial \omega / \partial E,$$



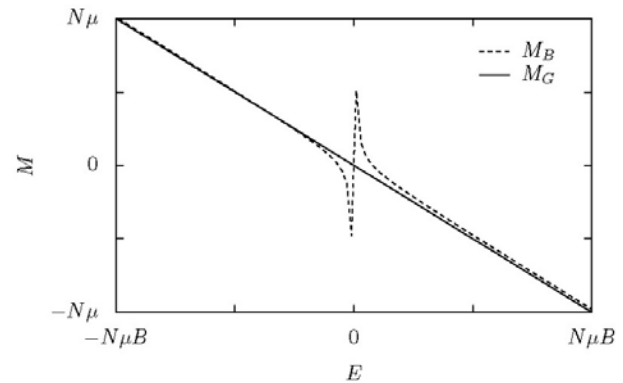
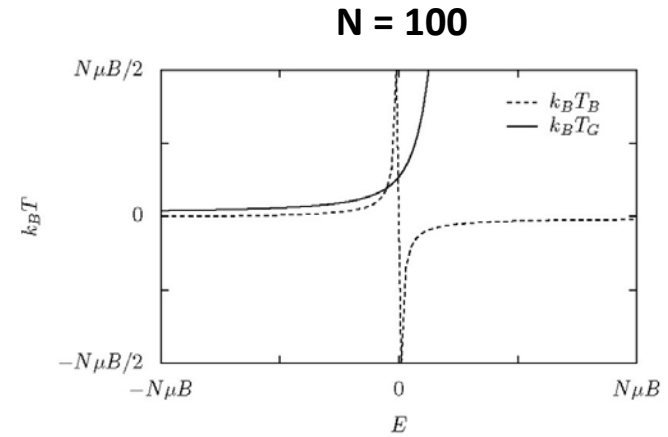
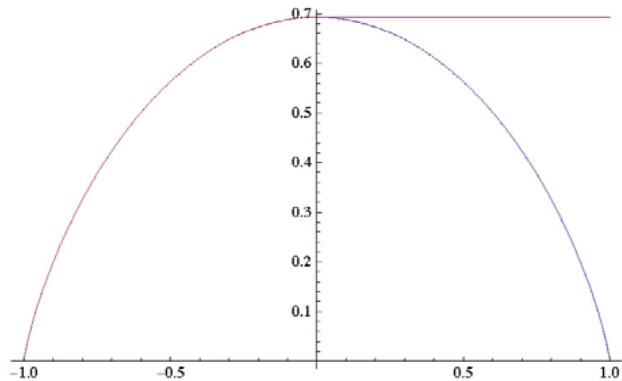
**Density of states of the pendulum in reduced units (complete elliptic integrals of the first kind).  
Fig. 1 in reference: M. Baeten and J. Naudts, *Entropy*, 13, 1186-1199 (2011).**

# N Spins $|\vec{S}| = 1/2$

Entropy for  $N = 100$   
(magenta:  $S_G$  ; blue:  $S_B$ )



$N = 10^8$



$$\Delta = M_B - M = -k_B T_B / B$$

# Negative Absolute Temperature for Motional Degrees of Freedom

S. Braun,<sup>1,2</sup> J. P. Ronzheimer,<sup>1,2</sup> M. Schreiber,<sup>1,2</sup> S. S. Hodgman,<sup>1,2</sup> T. Rom,<sup>1,2</sup>  
I. Bloch,<sup>1,2</sup> U. Schneider<sup>1,2\*</sup>

Because negative temperature systems can absorb entropy while releasing energy, they give rise to several counterintuitive effects, such as Carnot engines with an efficiency greater than unity (4). Through a stability analysis for thermodynamic equilibrium, we showed that negative temperature states of motional degrees of freedom necessarily possess negative pressure (9) and are thus of fundamental interest to the description of dark energy in cosmology, where negative pressure is required to account for the accelerating expansion of the universe (10).

- ✓ Carnot efficiencies  $> 1$
- ✓ Dark Energy

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# Recent Experiments

SCIENCE VOL 339 4 JANUARY 2013

## Negative Absolute Temperature for Motional Degrees of Freedom

S. Braun,<sup>1,2</sup> J. P. Ronzheimer,<sup>1,2</sup> M. Schreiber,<sup>1,2</sup> S. S. Hodgman,<sup>1,2</sup> T. Rom,<sup>1,2</sup>  
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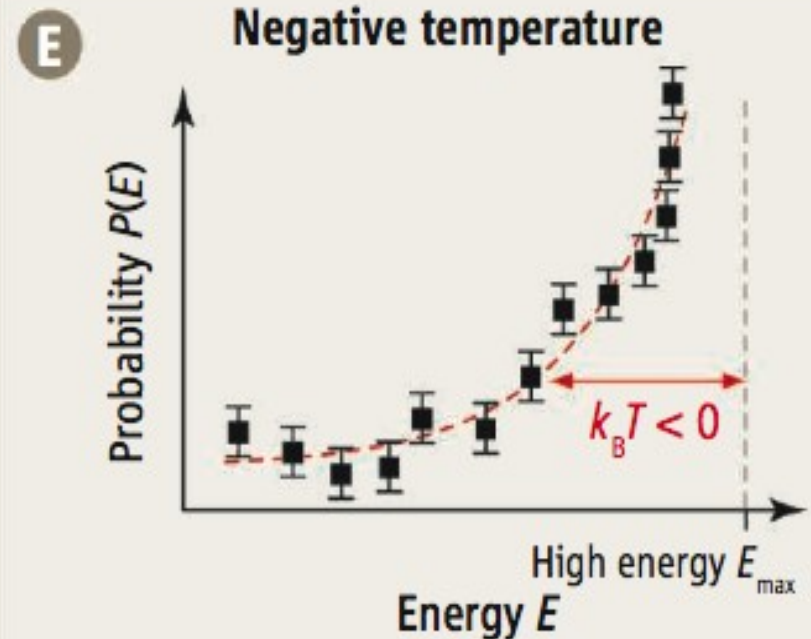
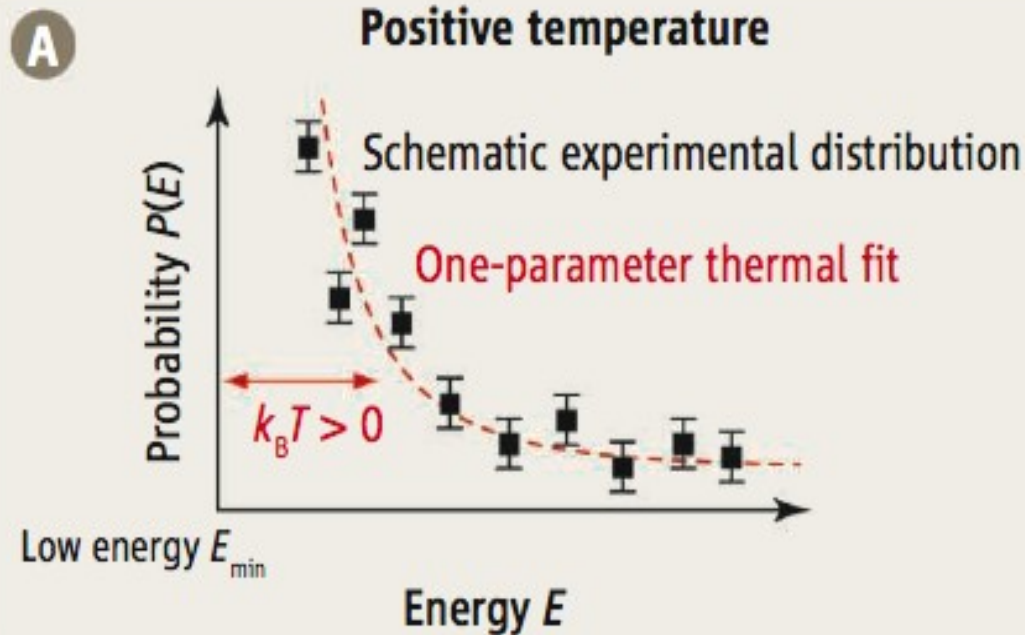
- fit to 1-particle level occupation  
⇒ **negative temperature**
  - Carnot efficiencies  $\eta > 1$
  - $T < 0$  &  $p < 0 \Rightarrow$  model for dark energy
-

# Recent Experiments

SCIENCE VOL 339 4 JANUARY 2013

## Negative Absolute Temperature for Motional Degrees of Freedom

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PHYSICAL REVIEW E **90**, 062116 (2014)



## Thermodynamic laws in isolated systems

Stefan Hilbert,<sup>1,\*</sup> Peter Hänggi,<sup>2,3</sup> and Jörn Dunkel<sup>4</sup>

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<sup>2</sup>*Institute of Physics, University of Augsburg, Universitätsstraße 1, D-86135 Augsburg, Germany*

<sup>3</sup>*Nanosystems Initiative Munich, Schellingstr. 4, D-80799 München, Germany*

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(Received 28 September 2014; published 9 December 2014)

**\*\* 23 pages \*\***

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## **Joint Work with**



**Jörn Dunkel**

**Department of Mathematics,  
Massachusetts Institute of Technology**

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**and**



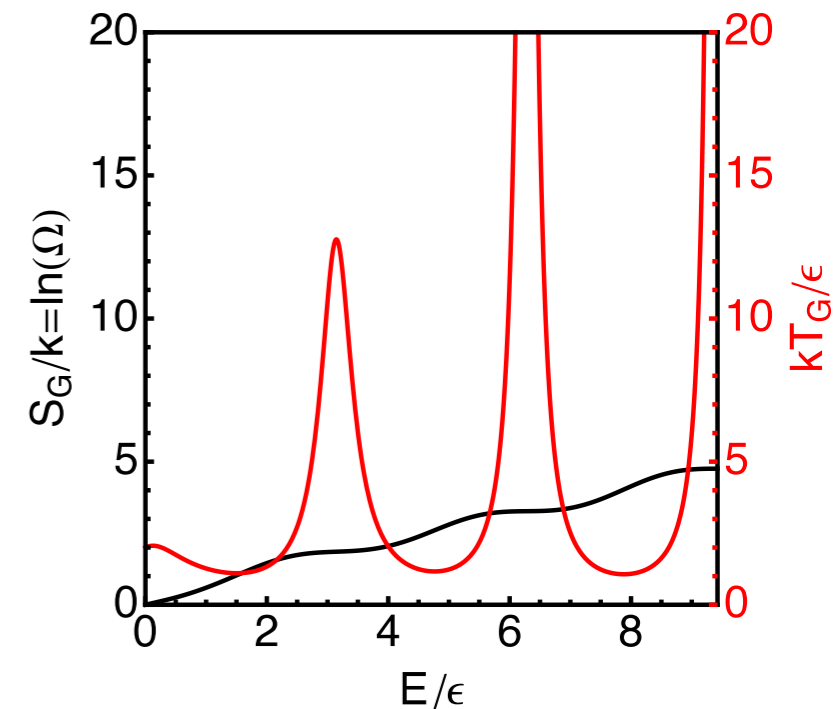
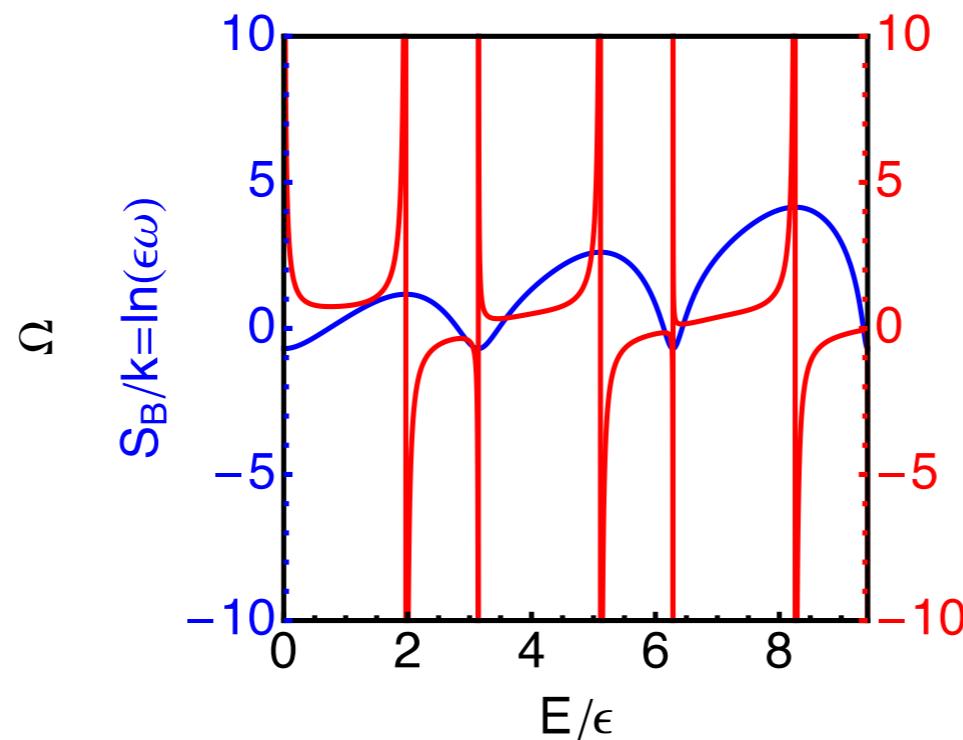
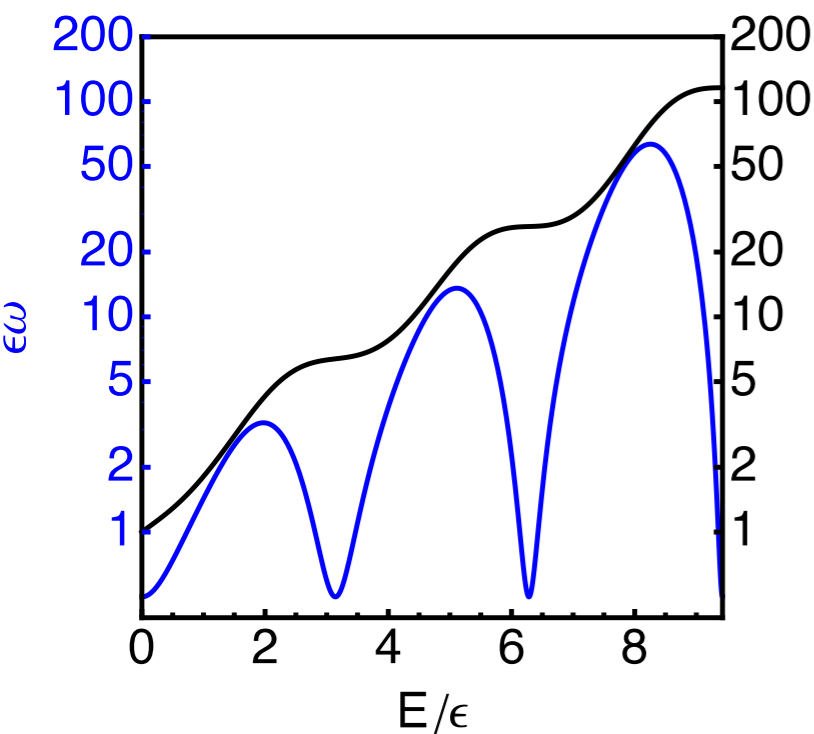
**Stefan Hilbert**

**Excellence Cluster Universe,  
Ludwig-Maximilians-Universität, Munich**

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# 'Non-uniqueness' of temperature

$$\Omega(E) = \exp \left[ \frac{E}{2\epsilon} - \frac{1}{4} \sin \left( \frac{2E}{\epsilon} \right) \right] + \frac{E}{2\epsilon},$$



Temperature does **NOT** determine direction heat flow.  
**Energy** is primary control parameter of MCE.

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# Thermodynamic Entropy

- $dE = \delta Q + \delta W$
- $W$  and  $Q$  **not** state functions  
 $\Rightarrow$  differentials  $\delta W$  and  $\delta Q$  not total
- but  $dS = \frac{\delta Q}{T}$  total differential

**if**  $\frac{1}{T} = \left( \frac{\partial f[\Omega(E, Z)]}{\partial E} \right)_Z$  **!**

---

# Mech. Adiabatic Processes

$$dE = \delta Q + \delta W = T dS - \sum_n p_n dZ_n$$

$$p_i = T \left( \frac{\partial S}{\partial Z_i} \right)_{E, Z_{n \neq i}}$$

$$dE = \sum_n \left\langle \frac{\partial H}{\partial Z_n} \right\rangle_{\rho} dZ_n$$

---

# Mech. Adiab. = Thermod. Adiab.

$$dE = \delta Q + \delta W = T dS - \sum_n p_n dZ_n$$

$$p_i = T \left( \frac{\partial S}{\partial Z_i} \right)_{E, Z_{n \neq i}}$$

$$dE = \sum_n \left\langle \frac{\partial H}{\partial Z_n} \right\rangle_{\rho} dZ_n$$

$$dS = 0$$

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# First Law

$$dE = \delta Q + \delta W = T dS - \sum_n p_n dZ_n$$

$$\Rightarrow p_i = T \left( \frac{\partial S}{\partial Z_i} \right)_{E, Z_{n \neq i}} \stackrel{!}{=} - \left\langle \frac{\partial H}{\partial Z_i} \right\rangle_{\rho}$$

$$\Rightarrow S(E, \dots) \stackrel{!}{=} f[\Omega(E, \dots)]$$

$f(\cdot) = k_B \ln(\cdot)$  from additional constraints  
(e.g. gas thermometer)



---



# First law

$$dE = \delta Q + \delta A = T dS - \sum_n p_n dZ_n$$

$$p_j = T \left( \frac{\partial S}{\partial Z_j} \right)_{E, Z_n \neq Z_j} \stackrel{!}{=} - \left\langle \frac{\partial H}{\partial Z_j} \right\rangle_E$$

Gibbs   $\Rightarrow$  Boltzmann 

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# Second Law

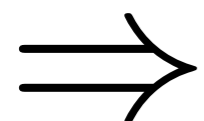
$$\sum_i^{\text{after}} S_i \stackrel{!}{\geq} \sum_j^{\text{before}} S_j$$

# Second law

Gibbs

$$S_G(E) = \ln \Omega$$

$$\begin{aligned}
 & \Omega(E_A + E_B) \\
 &= \int_0^{E_A + E_B} dE' \Omega_{\mathcal{A}}(E') \omega_{\mathcal{B}}(E_A + E_B - E') \\
 &= \int_0^{E_A + E_B} dE' \int_0^{E'} dE'' \omega_{\mathcal{A}}(E'') \omega_{\mathcal{B}}(E_A + E_B - E') \\
 &\geq \int_{E_A}^{E_A + E_B} dE' \int_0^{E_A} dE'' \omega_{\mathcal{A}}(E'') \omega_{\mathcal{B}}(E_A + E_B - E') \\
 &= \int_0^{E_A} dE'' \omega_{\mathcal{A}}(E'') \int_0^{E_B} dE''' \omega_{\mathcal{B}}(E''') \\
 &= \Omega_{\mathcal{A}}(E_A) \Omega_{\mathcal{B}}(E_B).
 \end{aligned}$$



$$S_{G_{\mathcal{A}\mathcal{B}}}(E_A + E_B) \geq S_{G_{\mathcal{A}}}(E_A) + S_{G_{\mathcal{B}}}(E_B)$$

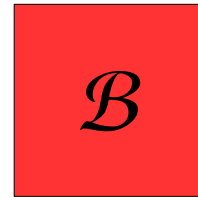
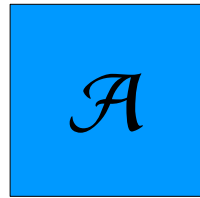


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# Second Law

before  
coupling

$$H_{\mathcal{A}} = E_{\mathcal{A}} \quad H_{\mathcal{B}} = E_{\mathcal{B}}$$

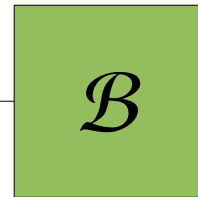


$$S_{\mathcal{A}}(E_{\mathcal{A}})$$

$$S_{\mathcal{B}}(E_{\mathcal{B}})$$

after  
coupling

$$H_{\mathcal{AB}} = H_{\mathcal{A}} + H_{\mathcal{B}} = E_{\mathcal{A}} + E_{\mathcal{B}} = E_{\mathcal{AB}}$$

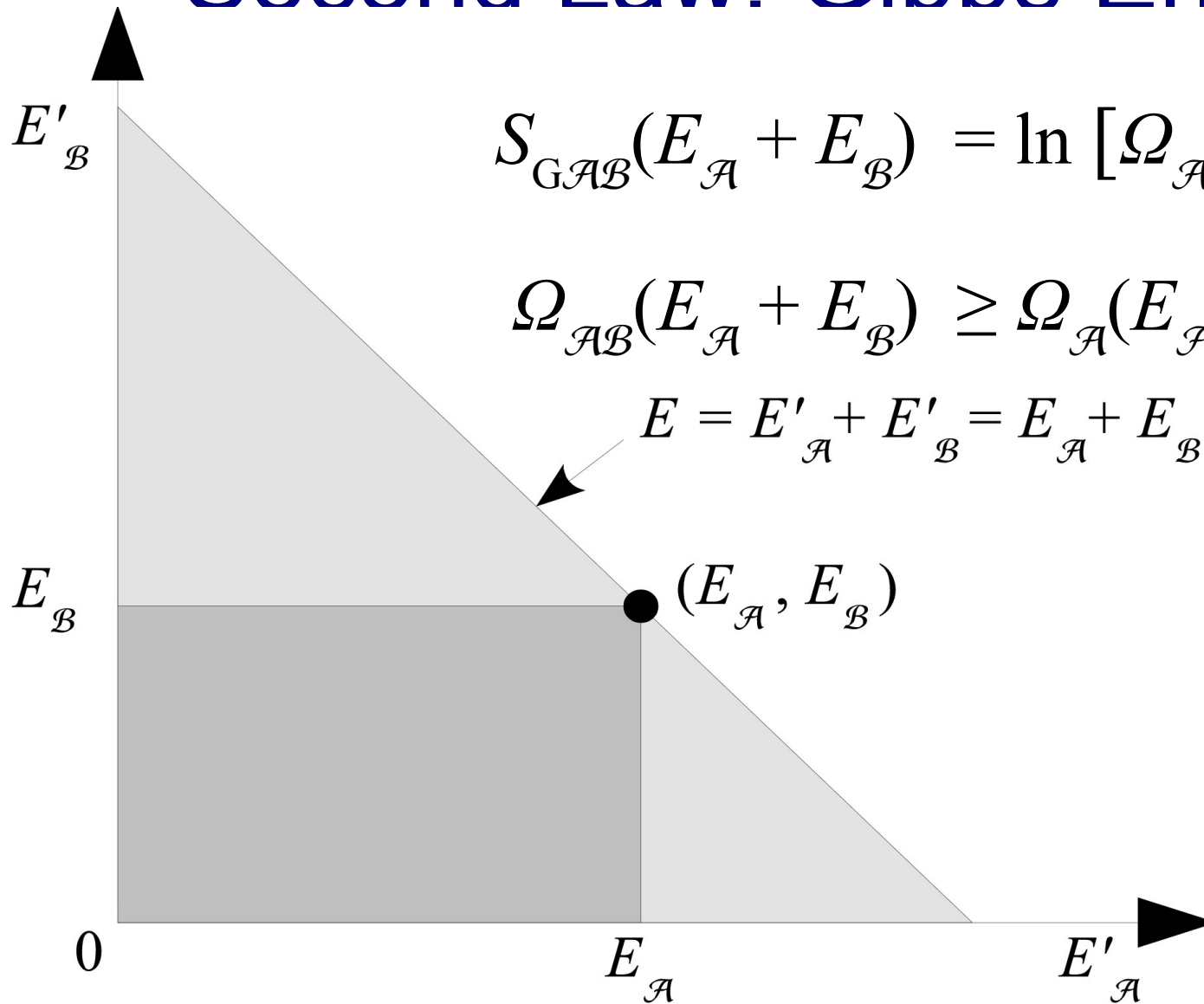


!

$$S_{\mathcal{AB}}(E_{\mathcal{AB}}) \geq S_{\mathcal{A}}(E_{\mathcal{A}}) + S_{\mathcal{B}}(E_{\mathcal{B}})$$

---

# Second Law: Gibbs Entropy



$$S_{GAB}(E_{\mathcal{A}} + E_{\mathcal{B}}) = \ln [\Omega_{AB}(E_{\mathcal{A}} + E_{\mathcal{B}})]$$

$$\Omega_{AB}(E_{\mathcal{A}} + E_{\mathcal{B}}) \geq \Omega_{\mathcal{A}}(E_{\mathcal{A}}) \Omega_{\mathcal{B}}(E_{\mathcal{B}})$$

$$E = E'_{\mathcal{A}} + E'_{\mathcal{B}} = E_{\mathcal{A}} + E_{\mathcal{B}}$$

# Second law

Boltzmann

$$S_{\mathcal{B}}(E) = \ln(\epsilon \omega)$$

$$\epsilon \omega(E_{\mathcal{A}} + E_{\mathcal{B}}) = \epsilon \int_0^{E_{\mathcal{A}} + E_{\mathcal{B}}} dE' \omega_{\mathcal{A}}(E') \omega_{\mathcal{B}}(E_{\mathcal{A}} + E_{\mathcal{B}} - E')$$

$$\neq \epsilon^2 \omega_{\mathcal{A}}(E_{\mathcal{A}}) \omega_{\mathcal{B}}(E_{\mathcal{B}})$$



---

# Second Law

$$\sum_i^{\text{after}} S_i \stackrel{!}{\geq} \sum_j^{\text{before}} S_j$$

$$S_{\mathcal{B}\mathcal{A}\mathcal{B}}(E_{\mathcal{A}} + E_{\mathcal{B}}) \not\geq$$

$$S_{\mathcal{B}\mathcal{A}}(E_{\mathcal{A}}) + S_{\mathcal{B}\mathcal{B}}(E_{\mathcal{B}})$$

$$S_{\mathcal{G}\mathcal{A}\mathcal{B}}(E_{\mathcal{A}} + E_{\mathcal{B}}) \geq$$

$$S_{\mathcal{G}\mathcal{A}}(E_{\mathcal{A}}) + S_{\mathcal{G}\mathcal{B}}(E_{\mathcal{B}})$$

---

**Erunt multi qui, postquam mea scripta legerint, non ad contemplandum utrum vera sint quae dixerim, mentem convertent, sed solum ad disquirendum quomodo, vel iure vel iniuria, rationes meas labefactare possent.**

**G. Galilei, *Opere* (Ed. Naz., vol. I, p. 412)**

**There will be many who, when they will have read my paper, will apply their mind, not to examining whether what I have said is true, but only to seeking how, by hook or by crook, they could demolish my arguments.**



---

# Example: Classical Ideal Gas

$$\Omega(E, V, N) = \alpha(D, N) V^N E^{DN/2}$$

$$S_B(E, \dots) = k_B \ln \epsilon \omega(E, \dots)$$

$$E = \left( \frac{DN}{2} - 1 \right) k_B T_B$$

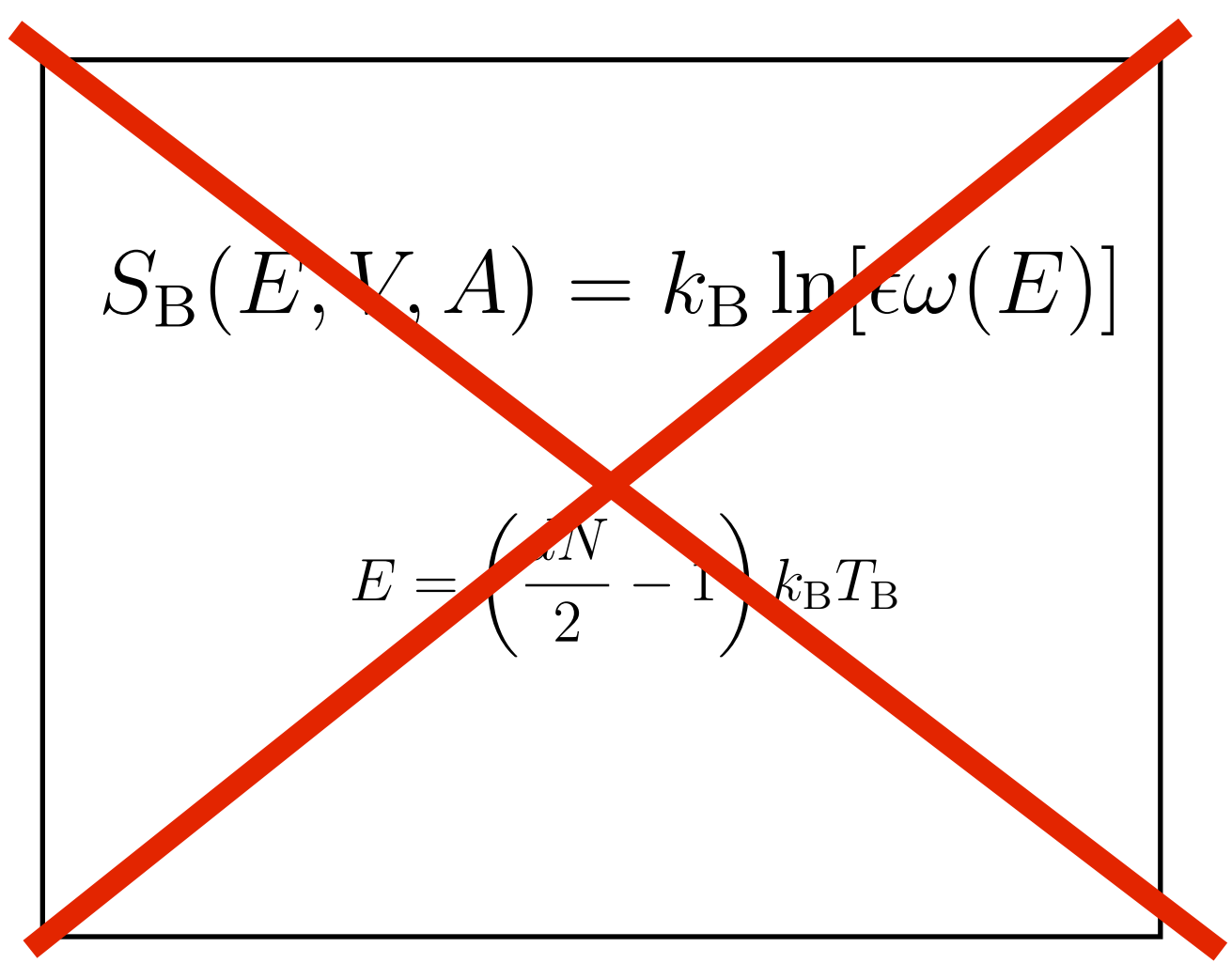
for  $DN = 1$  or  $DN = 2$ ?

$$S_G(E, \dots) = k_B \ln \Omega(E, \dots)$$

$$E = \frac{DN}{2} k_B T_G$$

# Example I: Classical ideal gas

$$\Omega(E, V) = \alpha E^{dN/2} V^N, \quad \alpha = \frac{(2\pi m)^{dN/2}}{N! h^d \Gamma(dN/2 + 1)}$$


$$S_B(E, V, A) = k_B \ln[\epsilon \omega(E)]$$

$$E = \left( \frac{dN}{2} - 1 \right) k_B T_B$$

**vs.**

$$S_G(E, V, A) = k_B \ln[\Omega(E)]$$

$$E = \frac{dN}{2} k_B T_G$$

---

# Example: Single 1-dim Particle

- single particle with energy  $E$   
in one-dimensional box of length  $L$
- $\Omega(E) \propto \sqrt{E}$

$$p_B = -\frac{2E}{L} \neq \left\langle \frac{\partial H}{\partial L} \right\rangle_\rho$$

(dark energy?)

$$p_G = +\frac{2E}{L} = \left\langle \frac{\partial H}{\partial L} \right\rangle_\rho$$

---

# Thermal equilibrium

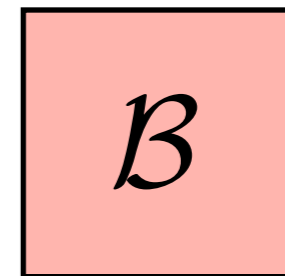
before  
coupling

$$H_A = E_A$$



$$T_A(E_A)$$

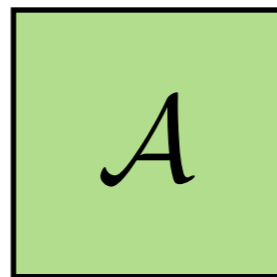
$$H_B = E_B$$



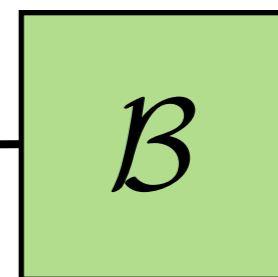
$$T_B(E_B)$$

after  
coupling

$$H = H_A + H_B = E_A + E_B = E$$



$$T(E)$$



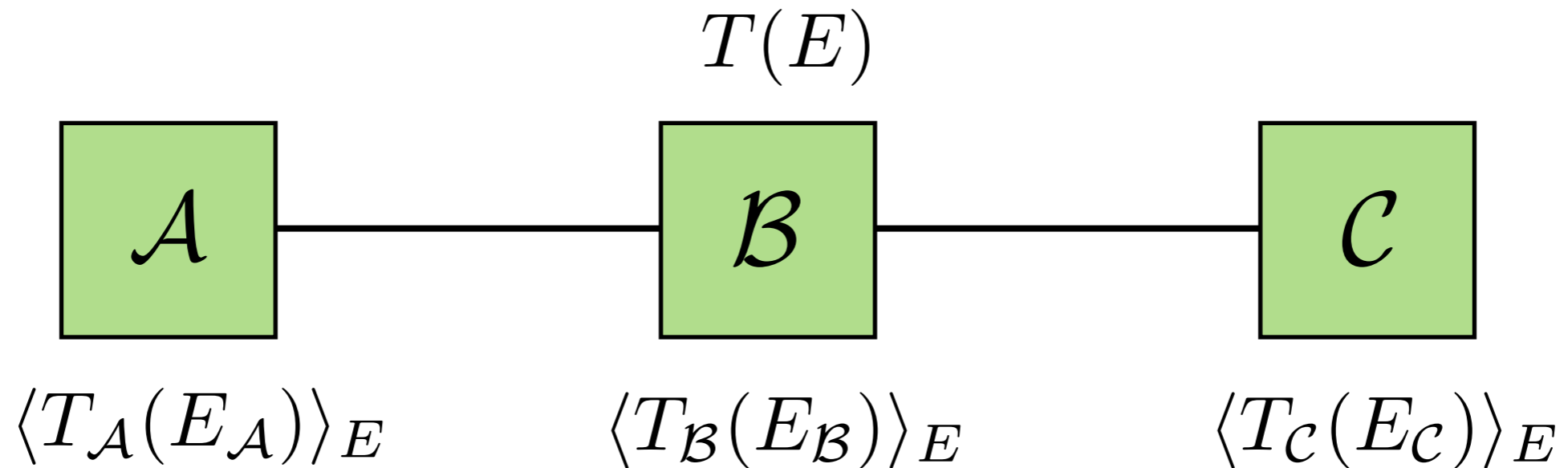
$$\langle T_A(E_A) \rangle_E$$

$$\langle T_B(E_B) \rangle_E$$

$$\langle T_i(E_i) \rangle_E = \int_0^\infty dE_i T_i(E_i) \pi_i(E_i|E)$$

$$\pi_A(E_A|E) = \frac{\omega_A(E_A) \omega_B(E - E_A)}{\omega(E)}$$

# Zeroth law



$$\langle T_i(E_i) \rangle_E \stackrel{!}{=} T(E)$$

## Gibbs

$$\begin{aligned}
 \langle T_{G_A}(E_A) \rangle_E &= \int_0^\infty dE_A \frac{\Omega_A(E_A)}{\omega_A(E_A)} \frac{\omega_A(E_A) \omega_B(E - E_A)}{\omega(E)} \\
 &= \frac{1}{\omega(E)} \int_0^E dE_A \Omega_A(E_A) \omega_B(E - E_A) \\
 &= T_G(E).
 \end{aligned}$$



---

# Summary

- Entropy candidates for isolated systems:

entropy	$S(E)$	0 <sup>th</sup> law	1 <sup>st</sup> law	2 <sup>nd</sup> law	equipart.
...other...					
Boltzmann	$\ln(\epsilon\omega)$	no	no	no	no
Gibbs	$\ln(\Omega)$	yes*	yes	yes	yes

canonical ensemble

$$\omega(E) = \text{Tr}[\delta(E - H)]$$

$$= \text{Tr}_{\mathcal{A}} \{ \text{Tr}_{\mathcal{B}} [\delta(E - H_{\mathcal{A}} - H_{\mathcal{B}})] \}$$

$$= \text{Tr}_{\mathcal{A}} \left\{ \text{Tr}_{\mathcal{B}} \left[ \int_{-\infty}^{\infty} dE'_{\mathcal{A}} \delta(E'_{\mathcal{A}} - H_{\mathcal{A}}) \int_{-\infty}^{\infty} dE'_{\mathcal{B}} \delta(E'_{\mathcal{B}} - H_{\mathcal{B}}) \delta(E - H_{\mathcal{A}} - H_{\mathcal{B}}) \right] \right\}$$

$$= \text{Tr}_{\mathcal{A}} \left\{ \text{Tr}_{\mathcal{B}} \left[ \int_{-\infty}^{\infty} dE'_{\mathcal{A}} \delta(E'_{\mathcal{A}} - H_{\mathcal{A}}) \int_{-\infty}^{\infty} dE'_{\mathcal{B}} \delta(E'_{\mathcal{B}} - H_{\mathcal{B}}) \delta(E - E'_{\mathcal{A}} - E'_{\mathcal{B}}) \right] \right\}$$

$$= \int_{-\infty}^{\infty} dE'_{\mathcal{A}} \text{Tr}_{\mathcal{A}}[\delta(E'_{\mathcal{A}} - H_{\mathcal{A}})] \int_{-\infty}^{\infty} dE'_{\mathcal{B}} \text{Tr}_{\mathcal{B}}[\delta(E'_{\mathcal{B}} - H_{\mathcal{B}})] \delta(E - E'_{\mathcal{A}} - E'_{\mathcal{B}})$$

$$= \int_{-\infty}^{\infty} dE'_{\mathcal{A}} \omega_{\mathcal{A}}(E'_{\mathcal{A}}) \int_{-\infty}^{\infty} dE'_{\mathcal{B}} \omega_{\mathcal{B}}(E'_{\mathcal{B}}) \delta(E - E'_{\mathcal{A}} - E'_{\mathcal{B}})$$

$$= \int_0^{\infty} dE'_{\mathcal{A}} \int_0^{\infty} dE'_{\mathcal{B}} \omega_{\mathcal{A}}(E'_{\mathcal{A}}) \omega_{\mathcal{B}}(E'_{\mathcal{B}}) \delta(E - E'_{\mathcal{A}} - E'_{\mathcal{B}})$$

$$= \int_0^E dE'_{\mathcal{A}} \omega_{\mathcal{A}}(E'_{\mathcal{A}}) \omega_{\mathcal{B}}(E - E'_{\mathcal{A}}).$$

$$\pi_{\mathcal{A}}(E'_{\mathcal{A}}|E) = \text{Tr}[\rho \delta(E'_{\mathcal{A}} - H_{\mathcal{A}})]$$

$$= \text{Tr} \left[ \frac{\delta(E - H_{\mathcal{A}} - H_{\mathcal{B}})}{\omega(E)} \delta(E'_{\mathcal{A}} - H_{\mathcal{A}}) \right]$$

$$= \int_{-\infty}^{\infty} dE''_{\mathcal{A}} \omega_{\mathcal{A}}(E''_{\mathcal{A}}) \int_{-\infty}^{\infty} dE''_{\mathcal{B}} \omega_{\mathcal{B}}(E''_{\mathcal{B}}) \frac{\delta(E - E''_{\mathcal{A}} - E''_{\mathcal{B}})}{\omega(E)} \delta(E'_{\mathcal{A}} - E''_{\mathcal{A}})$$

$$= \frac{\omega_{\mathcal{A}}(E'_{\mathcal{A}}) \omega_{\mathcal{B}}(E - E'_{\mathcal{A}})}{\omega(E)}.$$



# canonical ensemble

$$S^T = \delta(E^T - H(\xi, Z)) / \omega^T(E^T, Z) \Rightarrow P(E^S | E^T, Z) = \frac{\omega^S(E^S) \omega^B(E^T - E^S)}{\omega^T(E^T)}$$

$$E^T = E^S + E^B$$

$$= \frac{\omega^S(E^S)}{\mathcal{E} \omega^T(E^T)} \exp \left[ \frac{S_B^B(E^T - E^S)}{k_B} \right]$$

NEXT:  $S_B^B(E^T - E^S) = S_B^B(\bar{E}^B) + \frac{1}{T_B^B(\bar{E}^B)} (E^T - E^S - \bar{E}^B) + \dots,$

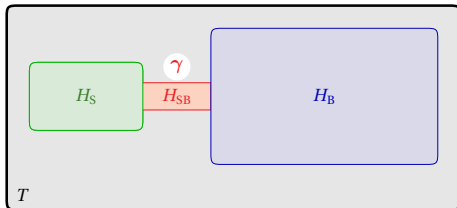
$$\Rightarrow \frac{\omega^S(E^S)}{\mathcal{E} \omega^T(E^T)} \exp \left[ \frac{S_B^B(\bar{E}^B)}{k_B} + \frac{(E^T - \bar{E}^B) - E^S}{k_B T_B^B(\bar{E}^B)} + \dots \right]$$

with  $+\dots \rightarrow 0$ !  $(\partial^2 S_B^B / \partial^2 E^B) = -1/T_B^2 C_B^B$



$$P(E^S | E^T, Z) = \frac{\omega^S(E^S)}{\mathcal{Z}_{can}} \exp \left[ - \frac{E^S}{k_B T_B^B(\bar{E}^B)} \right]$$

note:  $T_B^B(\bar{E}^B) \stackrel{?}{=} T_B^B(E^T)$ ; IF "normal":  $T_B^B = T_G^B = T_G^S = T_G^T$



The definition of thermodynamic quantities for systems coupled to a bath with finite coupling strength is not unique.

P. Hänggi, GLI, Acta Phys. Pol. B **37**, 1537 (2006)

# Partition function

$$Z_S(t) = \frac{Y(t)}{Z_B}$$

where  $Z_B = \text{Tr}_B e^{-\beta H_B}$

# An important difference



Route I

$$E \doteq E_S = \langle H_S \rangle = \frac{\text{Tr}_{S+B}(H_S e^{-\beta H})}{\text{Tr}_{S+B}(e^{-\beta H})}$$

Route II

$$\mathcal{Z} = \frac{\text{Tr}_{S+B}(e^{-\beta H})}{\text{Tr}_B(e^{-\beta H_B})} \quad U = -\frac{\partial \ln \mathcal{Z}}{\partial \beta}$$

$$\begin{aligned} \Rightarrow U &= \langle H \rangle - \langle H_B \rangle_B \\ &= E_S + \left[ \langle H_{SB} \rangle + \langle H_B \rangle - \langle H_B \rangle_B \right] \end{aligned}$$

For finite coupling  $E$  and  $U$  differ!

Quantum  
Brownian  
motion and  
the 3<sup>rd</sup> law

Specific heat and  
dissipation

Two approaches  
Microscopic model

Route I

Route II  
specific heat  
density of states

Conclusions



## Strong coupling: Example

System: Two-level atom; "bath": Harmonic oscillator

$$H = \frac{\epsilon}{2} \sigma_z + \Omega \left( a^\dagger a + \frac{1}{2} \right) + \chi \sigma_z \left( a^\dagger a + \frac{1}{2} \right)$$

$$H^* = \frac{\epsilon^*}{2} \sigma_z + \gamma$$

$$\epsilon^* = \epsilon + \chi + \frac{2}{\beta} \operatorname{artanh} \left( \frac{e^{-\beta\Omega} \sinh(\beta\chi)}{1 - e^{-\beta\Omega} \cosh(\beta\chi)} \right)$$

$$\gamma = \frac{1}{2\beta} \ln \left( \frac{1 - 2e^{-\beta\Omega} \cosh(\beta\chi) + e^{-2\beta\Omega}}{(1 - e^{-\beta\Omega})^2} \right)$$

$$Z_S = \operatorname{Tr} e^{-\beta H^*} \quad F_S = -k_b T \ln Z_S$$

$$S_S = -\frac{\partial F_S}{\partial T} \quad C_S = T \frac{\partial S_S}{\partial T}$$

# Free energy of a system strongly coupled to an environment

Thermodynamic argument:

$$F_S = F - F_B^0$$

$F$  total system free energy

$F_B$  bare bath free energy.

With this form of free energy the three laws of thermodynamics are fulfilled.

G.W. Ford, J.T. Lewis, R.F. O'Connell, Phys. Rev. Lett. **55**, 2273 (1985);

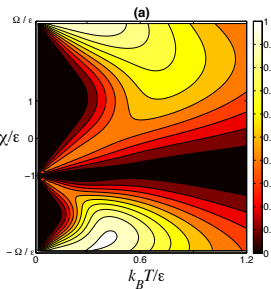
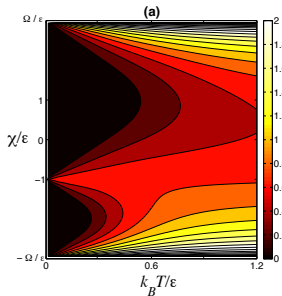
P. Hänggi, G.L. Ingold, P. Talkner, New J. Phys. **10**,115008 (2008);

G.L. Ingold, P. Hänggi, P. Talkner, Phys. Rev. E **79**, 0611505 (2009).

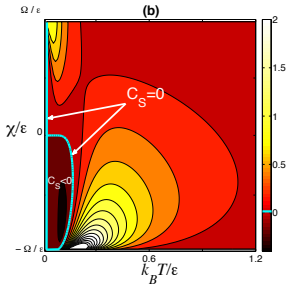
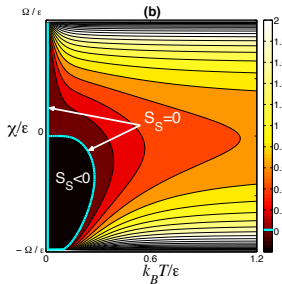
# Entropy and specific heat

Fluctuation  
Theorem for  
Arbitrary  
Open  
Quantum  
Systems

Michele  
Campisi



$$\Omega/\epsilon = 3$$



$$\Omega/\epsilon = 1/3$$

---

# Summary

## Isolated systems:

- thermodynamic entropy: Gibbs volume entropy
  - **no** negative thermodynamic temperature
  - **no** Carnot efficiencies  $> 1$
-



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# Summary

Isolated systems:

- T's before coupling do **not** predict heat flow during coupling
  - bounded spectra  $\Rightarrow$  ensembles **not** equivalent
  - thermodynamic entropy  
**not** always Shannon-like entropy  $S_S = -\text{Tr}[\rho \ln \rho]$
  - incorrect entropy  
 $\Rightarrow$  incorrect temperature, **forces**, and **responses**
-

---

# Summary

## Systems coupled to heat bath:

- Boltzmann temperature (of the bath) describes the probability of (energy)-fluctuations
- Boltzmann factor approximate for bath with finite heat capacity
- for certain  $N$ -particle isolated systems:  
1-p.  $p_i(E_i) \approx c \exp[-E_i / (k_B T_{B, N-1})]$

# PHILOSOPHICAL TRANSACTIONS A

[rsta.royalsocietypublishing.org](http://rsta.royalsocietypublishing.org)

Research



Article submitted to journal

**Subject Areas:**

Thermodynamic temperature

**Keywords:**

Thermodynamic ensembles, entropy, isolated systems, weak and strong coupling

## Meaning of temperature in different thermostistical ensembles

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Peter Hänggi<sup>1,2</sup>, Stefan Hilbert<sup>3</sup> and Jörn Dunkel<sup>4</sup>

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Important **UNSOLVED** (open) Problems are:

1.) **Quantum systems** and **discrete spectral parts**: DoS becomes singular  
==> a sum of delta-functions !!!

??? !!! best smoothing procedure ???!!!

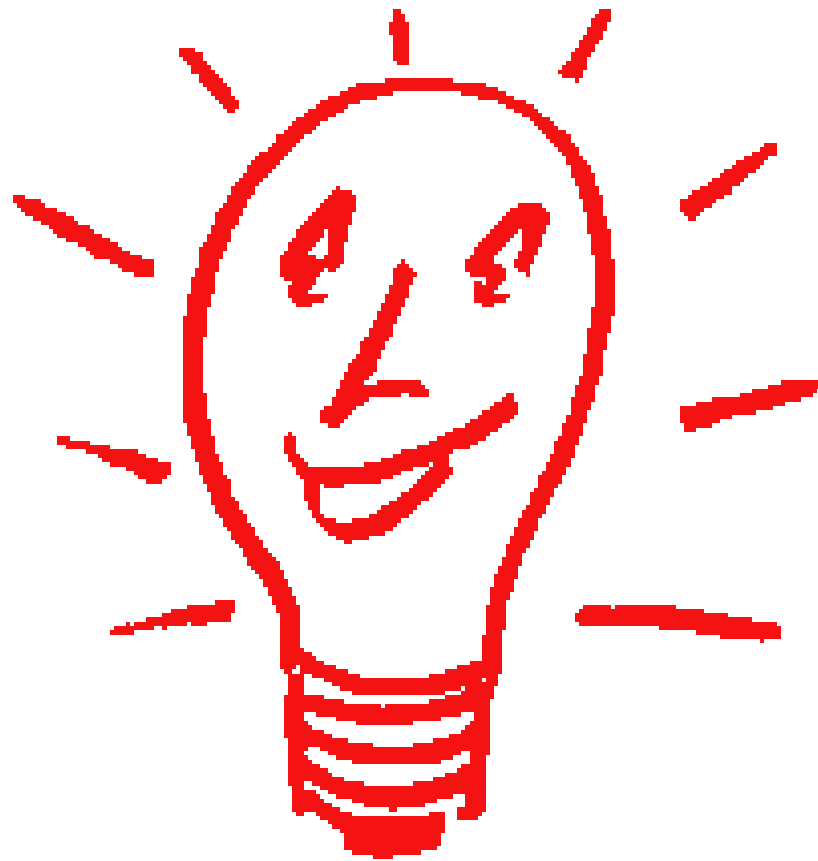
2.) **Canonical ensemble**: When is the Boltzmann factor truly OK?

3.) **Canonical ensemble and STRONG coupling**:

**Quantum case**: Canonical specific heat can now become **negative (!)**  
despite system being stable

**Classical case**: Are **\*negative\*** canonical specific heat values possible?

# A QUESTION ?



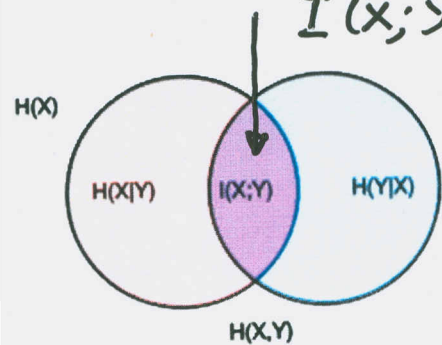
**Erunt multi qui, postquam mea scripta legerint, non ad contemplandum utrum vera sint quae dixerim, mentem convertent, sed solum ad disquirendum quomodo, vel iure vel iniuria, rationes meas labefactare possent.**

**G. Galilei, *Opere* (Ed. Naz., vol. I, p. 412)**

**There will be many who, when they will have read my paper, will apply their mind, not to examining whether what I have said is true, but only to seeking how, by hook or by crook, they could demolish my arguments.**

# Conditional Entropy

$$I(x; y) = H(x, y)$$



$$H(X) = H(X|Y) + I(X; Y)$$

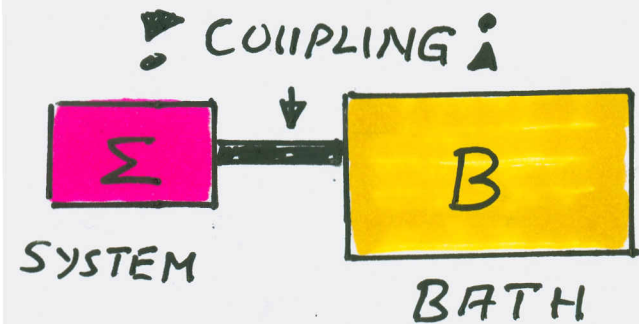
$$H(Y) = H(Y|X) + I(X; Y)$$

$$H(Y|X) = -\sum_{x,y} p(x,y) \ln p(x,y) - \left( -\sum_x p(x) \ln p(x) \right)$$

$$= -\sum_{x,y} p(x,y) \ln p(y|x) \geq 0$$

# Quantum Conditional Entropy

von NEUMANN:  $S_{vN}(\rho_\Sigma) = -\text{Tr} \rho_\Sigma \ln \rho_\Sigma$



$$S_\Sigma = S_{vN}(\rho_{\Sigma \times B}^{can}) - S_{vN}(\rho_B^{can})$$

can. thermod. entropy

quantum cond. entropy

$$\stackrel{?!}{=} S_{vN}(\rho_{\Sigma \times B}^{can}) - S_{vN}(\rho_B = \text{Tr}_\Sigma \rho_{\Sigma \times B}^{can})$$

$\uparrow \neq \rho_B^{can}$

$$T_B = \frac{T_G}{1 - k_B/C}$$

**Proof:**

$$k_B T_G = \frac{\Omega}{\Omega'} = \frac{\Omega}{\omega}, \quad k_B T_B = \frac{\omega}{\omega'} = \frac{\Omega'}{\Omega''}$$

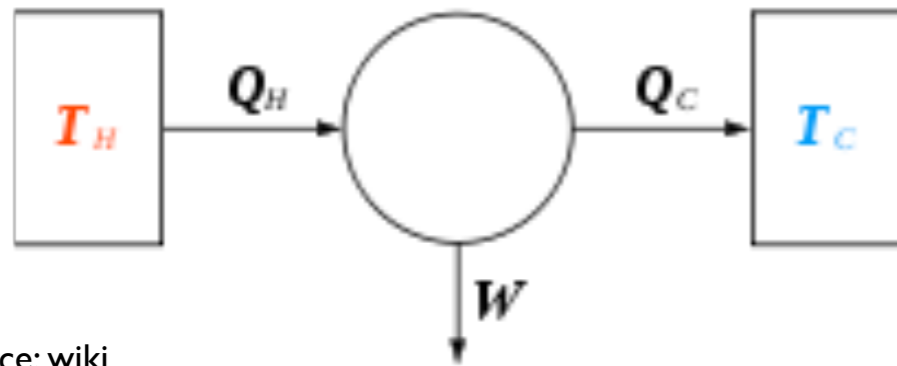
$$\begin{aligned} \frac{1}{C} &\equiv \left( \frac{\partial T_G}{\partial E} \right) = \frac{1}{k_B} \left( \frac{\Omega}{\Omega'} \right)' = \frac{1}{k_B} \frac{\Omega' \Omega' - \Omega \Omega''}{(\Omega')^2} \\ &= \frac{1}{k_B} \left[ 1 - \frac{\Omega \Omega''}{(\Omega')^2} \right] = \frac{1}{k_B} \left( 1 - \frac{T_G}{T_B} \right) \end{aligned}$$

□

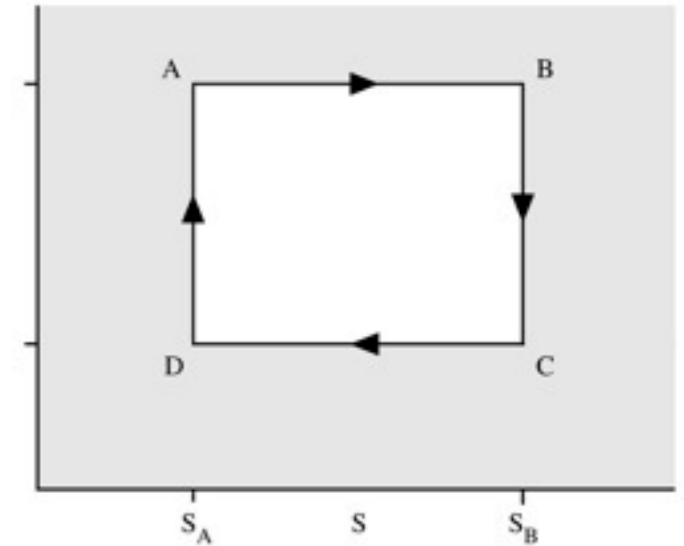
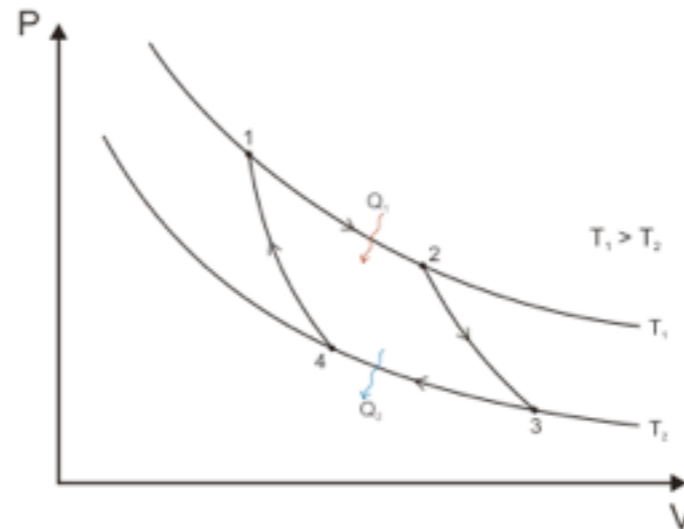


# Carnot efficiencies $> 1$ ?

$$\eta = 1 - \frac{T^C}{T^H}$$



source: wiki



- Carnot cycle assumes adiabatic switching
- Carnot formula assumes that TD relations are fulfilled .. only for  $T_G > 0$  the case

$\Rightarrow$  no Carnot efficiencies  $> 1$  when treated consistently